Development of a Full-Multigrid Gravity Solver for Athena++

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Features of Athena++

Already available	Not public yet	Being implemented	Planned
HD/MHD Curvilinear coordinate	Non-ideal MHD (Ohmic, Hall, AD) Radiation transfer	Particles (star / dust) Self-gravity (Multigrid on SMR/AMR)	General EOS Post-processing radiation transfer
SMR/AMR Special Relativity General Relativity	(Direct ray tracing) Self-gravity (FFT on uniform grid)	Heterogeneous parallelization	(ALMA Science Proj.) Hybrid PIC Plasma
(Fixed metric) MPI + OpenMP	Chemical reactions Shearing Box	Developer's guide Code paper	Radiation transfer (VTEF+implicit etc.) Full General Relativity
Parallel IO (MPI/HDF) User-defined functions	4th-order scheme Self-gravity (Multigrid		(dynamic metric)
Support for Intel, GCC, IBM, Cray, incl. KNL Website / tutorial	on uniform grid)		

赤字: Features (being) developed at Osaka

青字: Osaka is involved

Development of Multigrid Solver

Some physics requires solutions to global/implicit PDE.

- Self-gravity
- Radiation transfer (FLD, Variable Eddington Factor)
- Conduction, Diffusion, Viscosity

We need a good parallel elliptic/parabolic PDE solver

- fast and scalable
- robust and accurate
- compatible with AMR
- flexible and versatile → C++ derived classes

But how "fast" should it be?

→ at least faster than the MHD part.

Hyperbolic vs Ellipitical Eqs

Hydrodynamics:
$$\frac{\partial \boldsymbol{U}}{\partial t} + \nabla \cdot \boldsymbol{F} = 0$$

Complicated, but well established (Riemann solvers).

Information propagates at characteristic speeds

→ Finite Volume Method with local explicit update works well.

Gravity:

$$\nabla^2 \varphi = 4\pi G \rho$$

This is very simple, but one of the worst equations numerically.

- Information propagates instantaneously
- → Global, consistent solution is required
- Boundary conditions matter
- → Physically consistent boundaries are required
- Computationally cheap but memory / network intensive
- → Difficult to achieve good performance / scalability

Poisson Solvers

Discretized equation (2D, NxN cells):

$$\frac{\varphi_{j,i+1} - 2\varphi_{j,i} + \varphi_{j,i-1}}{\Delta x^2} + \frac{\varphi_{j+1,i} - 2\varphi_{j,i} + \varphi_{j-1,i}}{\Delta y^2} = 4\pi G \rho_{j,i}$$

$$A\varphi = \rho$$

It is hopeless to solve this equation directly: O(N⁶) (N⁹ in 3D).

Common numerical methods:

- Fast Fourier Transform (already implemented on Athena++)
 - + Deterministic (i.e. require only one sweep) and robust
 - + Computationally efficient
 - Efficient only on uniform and 2^N cells with periodic boundaries
 - Require global communication including global transpose
 - O(NlogN) logN is actually not a small factor
- Tree solver (e.g. Wünsh et al. 2017 for FLASH)
- Iterative solvers

Basic Iterative Solvers

$$A\boldsymbol{\varphi} = \boldsymbol{\rho} \rightarrow (L + D + U)\boldsymbol{\varphi} = \boldsymbol{\rho}$$

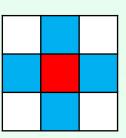
D = diagonal, L = lower triangle, U = upper triangle

Jacobi iteration: $\boldsymbol{\varphi}^{k+1} = D^{-1} \{ \boldsymbol{\rho} - (L+U) \boldsymbol{\varphi}^k \}$

Gauss-Seidel iteration: $\boldsymbol{\varphi}^{k+1} = D^{-1} \{ \boldsymbol{\rho} - L \boldsymbol{\varphi}^{k+1} - U \boldsymbol{\varphi}^k \}$

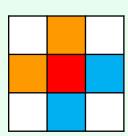
Jacobi:

Update a target cell using old information



Gauss-Seidel:

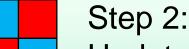
Use updated cells and old information



Improvement: Red-Black Gauss-Seidel iteration

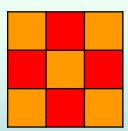
Step1:

Update **red cells** using **old** information



Update red cells using updated cells

⇒ Less dependent, isotropic scheme (but still slow)



Multigrid Concept

Iterative solvers = essentially **smoothing** (diffusion) of residual Diffusion "timescale" $\tau \sim \lambda^2/D \rightarrow$ small-scale noises diffuse faster \Rightarrow Accelerate convergence by applying different resolutions

2-level algorithm:

- 1. Coarsen (restrict) the grid to the coarse level
- 2. Apply diffusion (RBGS smoothing) on the coarse level
- 3. Project (prolongate) the result onto the fine level
- 4. Apply diffusion (RBGS smoothing) on the fine level
- 5. Repeat until the solution converges

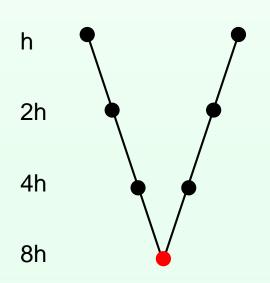
Configuration of our Multigrid:

V(1,1) Cycle = 1 smoothing before restriction, 1 after prolongation

Restriction: volume-weighted average

Prolongation: trilinear or triquadratic interpolation

Multigrid Poisson Solver



- Red-Black Gauss-Seidel Smoothing
- \ Restriction
- / Prolongation
- Solve Coarsest Level (exact or iterative)

Note: # of communications

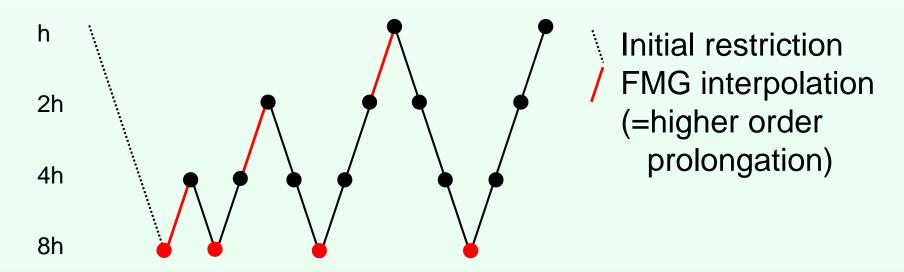
GSRB: 2 boundary communications

Restriction/Prolongation: 1 communication

V-Cycle MG solver can be applied as an iterative solver

- Reduces error in all the wavelengths
- Computationally efficient: O(N) (but O(NlogN) if communication dominates)
- Typically one sweep reduces the error by a factor of ~10
- Need a good initial guess to get fast convergence
- Memory and network intensive (especially latency as it requires a lot of small messages)

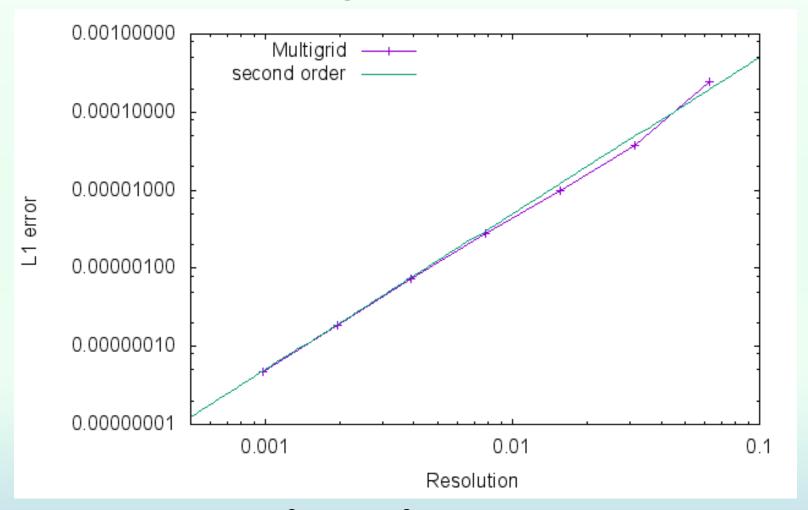
Full Multigrid Method



FMG solver: starting from the coarsest level, and use the result of the V-cycle MG solver as the initial guess on the finer grid.

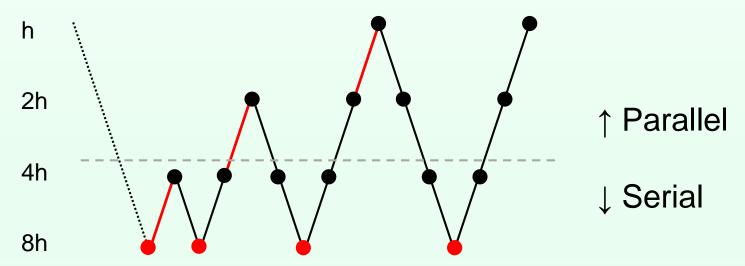
- One FMG sweep is sufficient to achieve ~ discretization error
- Extremely efficient: cost ~ 2 full V-cycles
- Higher order interpolation is needed (trilinear for normal prolongation, triquadratic for FMG)

FMG Convergence



3D sinusoidal waves, 16³→1024³ / wavelength Second-order convergence is achieved with one FMG sweep.

Parallel Multigrid

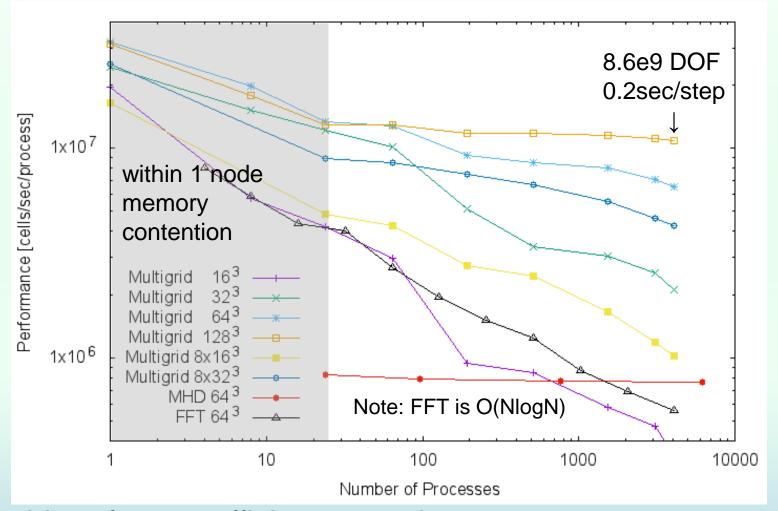


In Athena++, the domain is decomposed into MeshBlocks. At somepoint, a MeshBlocks is derefined into 1x1x1 cell → collect data using MPI_Allgather and apply Multigrid,

Gauss-Seidel smoother is cheap → network latency matters Using Athena++'s dynamic scheduling with TaskList, we interleave (overlap) communication and computation (This works only with more than one MeshBlocks/process)

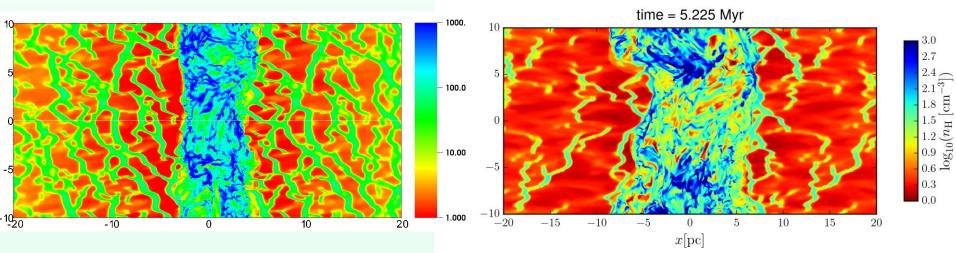


Cray XC30 @ NAOJ Flat MPI



Multigrid performs sufficiently good at least up to 4096 cores Small MeshBlocks are slow, but it is OK with many MeshBlocks

Molecular Cloud Formation



Left: with self-gravity Right: without self-gravity Caution: this is not a fair comparison - the density is twice higher. 1024x512x512, <n>=10 or 5 / cc, v=20km/s, $B=5\mu G$, $\theta=11^{\circ}$

After ~5Myr the maximum density reaches >10 6 /cc and collapses. The accumulated mass is so large that self-gravity is significant. \rightarrow to follow formation of cloud cores, we need SMR / AMR.

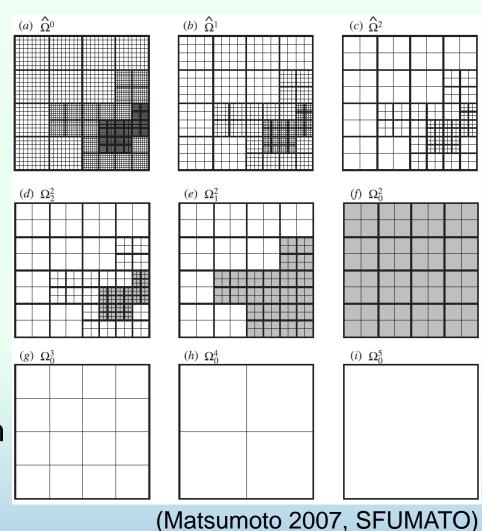
(see Iwasaki-san's talk for details) Only 20% additional cost!

Next Step: SMR/AMR

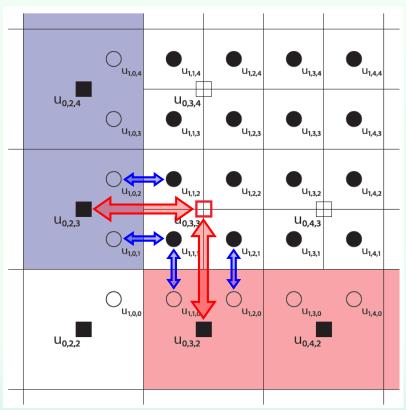
Basically, we will follow SFUMATO's approach→

On SMR/AMR grids, the Full Approximation Scheme (FAS) must be adopted.

Mass conservation formula (Feng et al., JCoPh, 2017, 352, 463) consistently discretize the level boundary so that the fluxes on the coarse and fine levels match → no need of correction, possibly improve convergence



Mass Conservation Formula



(Feng et al. JCoPh, 2017, 352, 463)

: Coarse active cells

: Fine active cells

O: Ghost cells

☐:Zombie cells

Ghost cells ○ are set by quadratic interpolation using active cells ■ ○. Zombie cells are calculated so that the fluxes between levels match.

Note that the double zombie cell has two different values – but it is OK as this cell appears only as a boundary value between the levels.

Summary

Multigrid works!

- Full Multi Grid method works very well
- Achieved sufficiently high performance
- Flexible interface using derived classes: can be applied to other physics such as radiation transfer
- Heterogeneous parallelization will improve performance

Athena++

- Planning next public release soon
- Non-ideal MHD, shearing box, chemistry, higher order etc...
- Even more physics: gravity, full GR, radiation, ...
- Method paper in progress