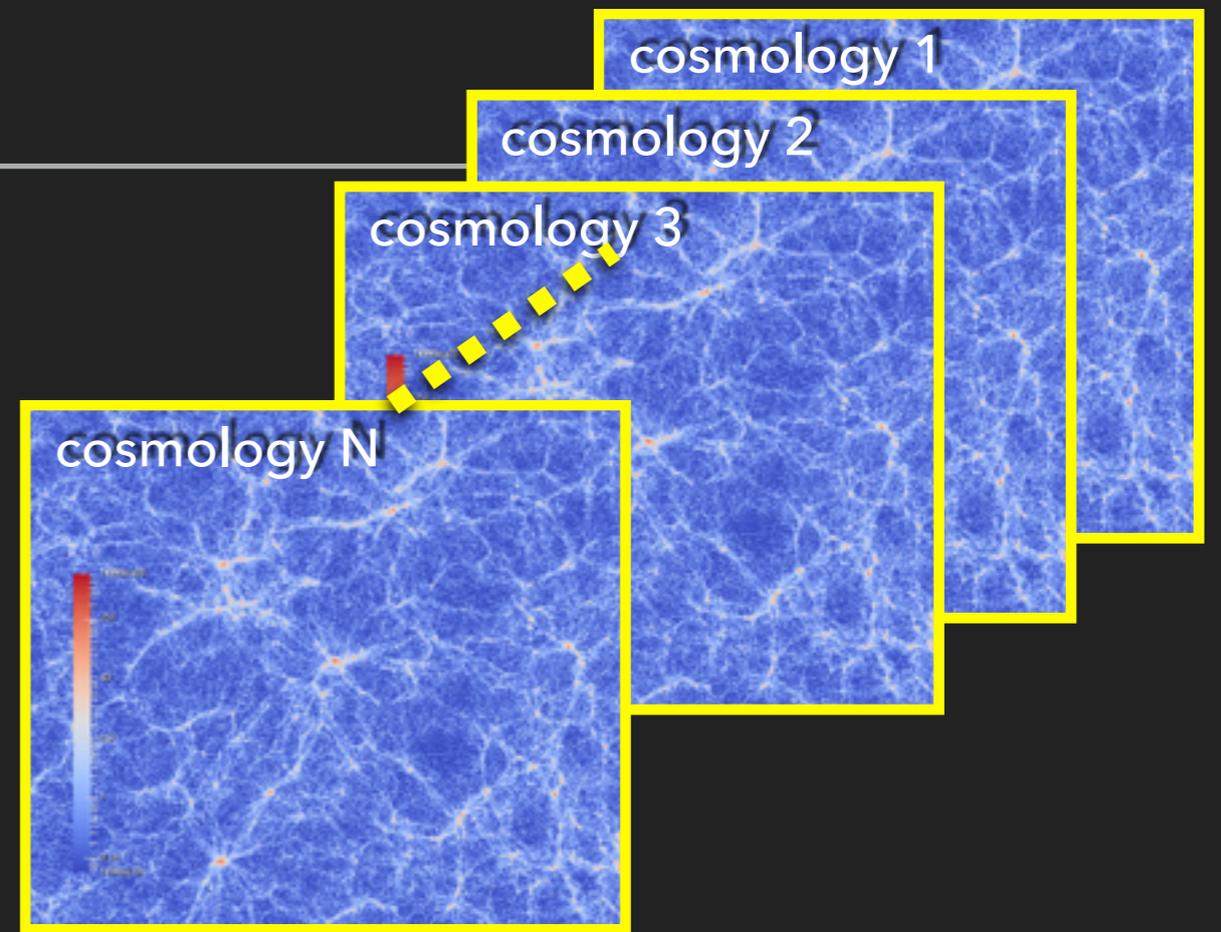
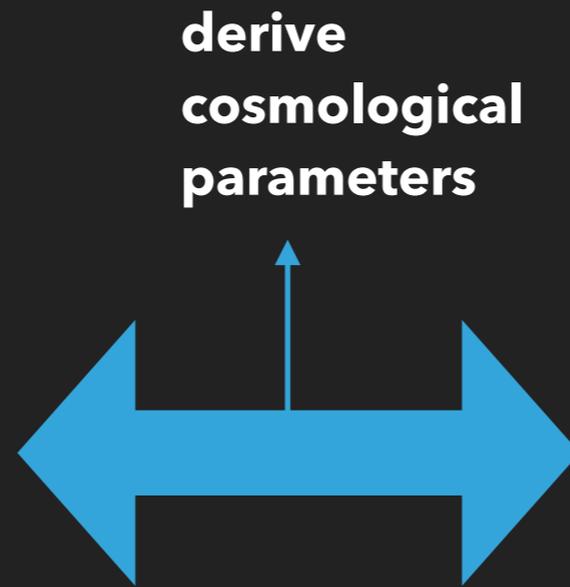
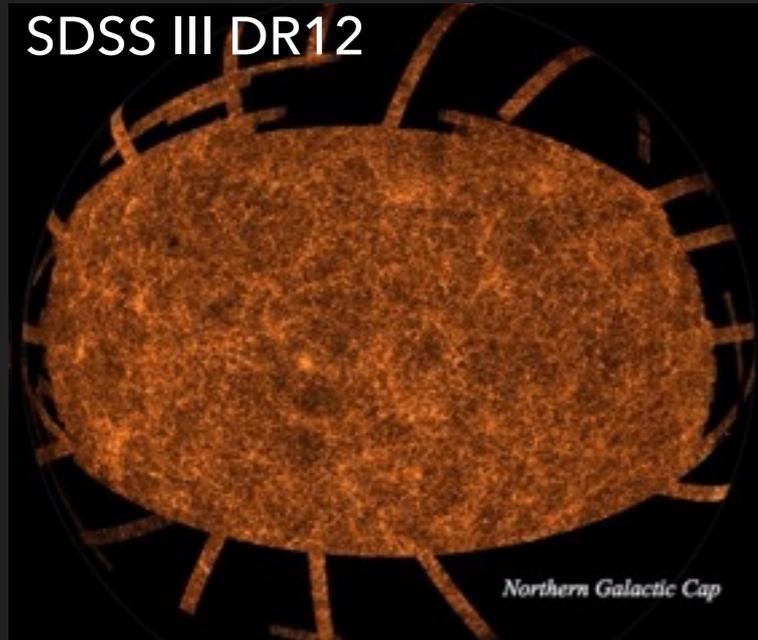


大規模宇宙論的シミュレーション群を用いた宇宙大規模構造の精密理論テンプレートの構築

TAKAHIRO NISHIMICHI (KAVLI IPMU, JST CREST)

THE ULTIMATE GOAL



Establishment of Numerical Cosmology

- ✓ Direct comparison btwn simulation and observation to draw cosmological information
- ✓ N-body method kind of established for mock large scale structures by solving gravitational system

- ✓ "Scan" multi-D cosmological parameter space?
 - ✓ feasibility questionable given the high dimensionality and required accuracy for future programs
 - ✓ **Efficient sampling scheme and use of machine learning techniques**
 - ✓ **Quick simulations with new algorithm**

GALAXY-GALAXY LENSING SIMULATION EFFORT

Simulations for Subaru HSC

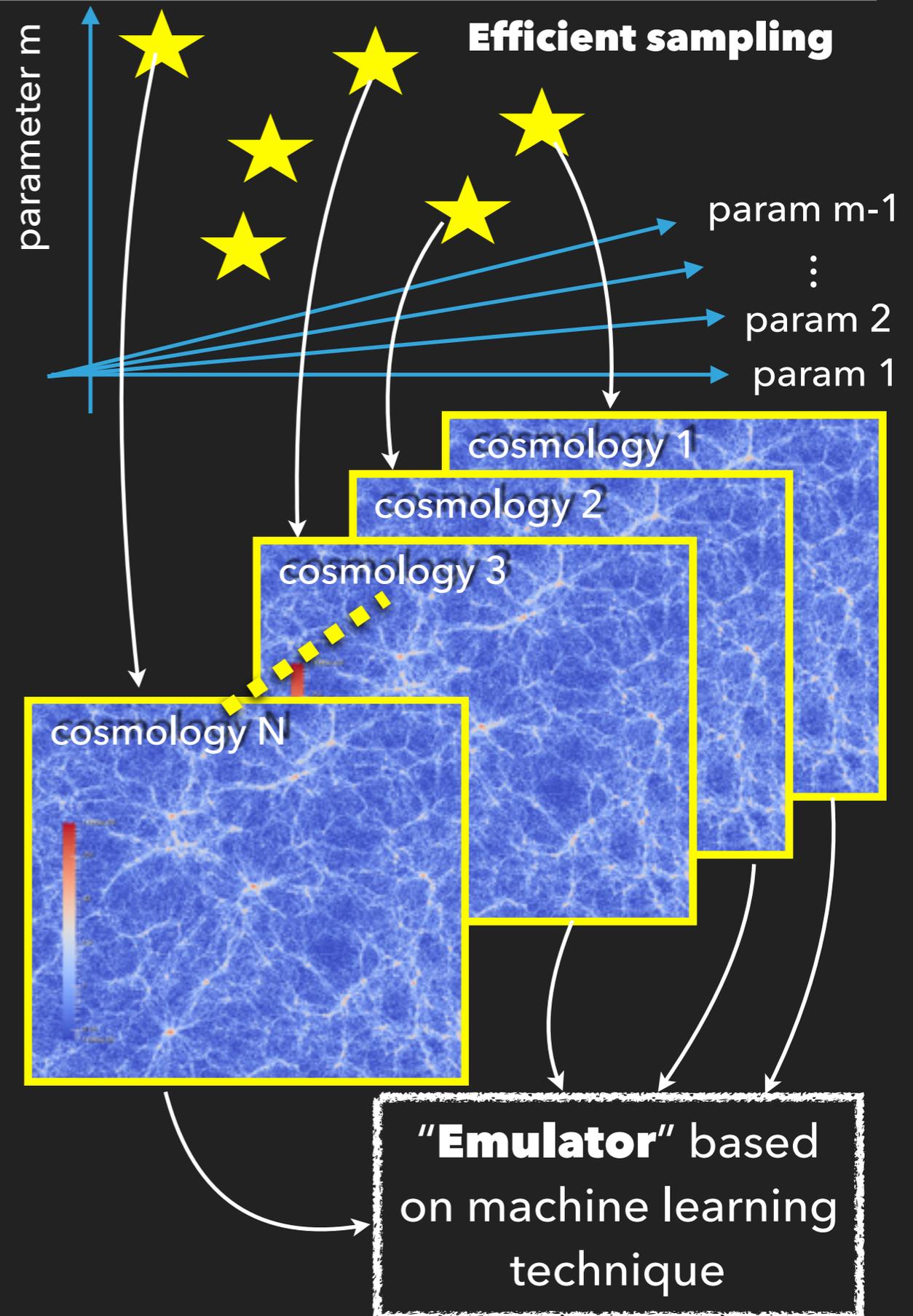
- ✓ accurate theoretical templates
 - ✓ Galaxy (cluster)-galaxy lensing
 - ✓ Galaxy-3D spatial clustering
- cross-correlation analysis w/ BOSS CMASS
- dark matter (sub)-halos as a simplest representative of galaxies

▶ **Kavli IPMU**
Takahiro Nishimichi*
Masahiro Takada
Naoki Yoshida

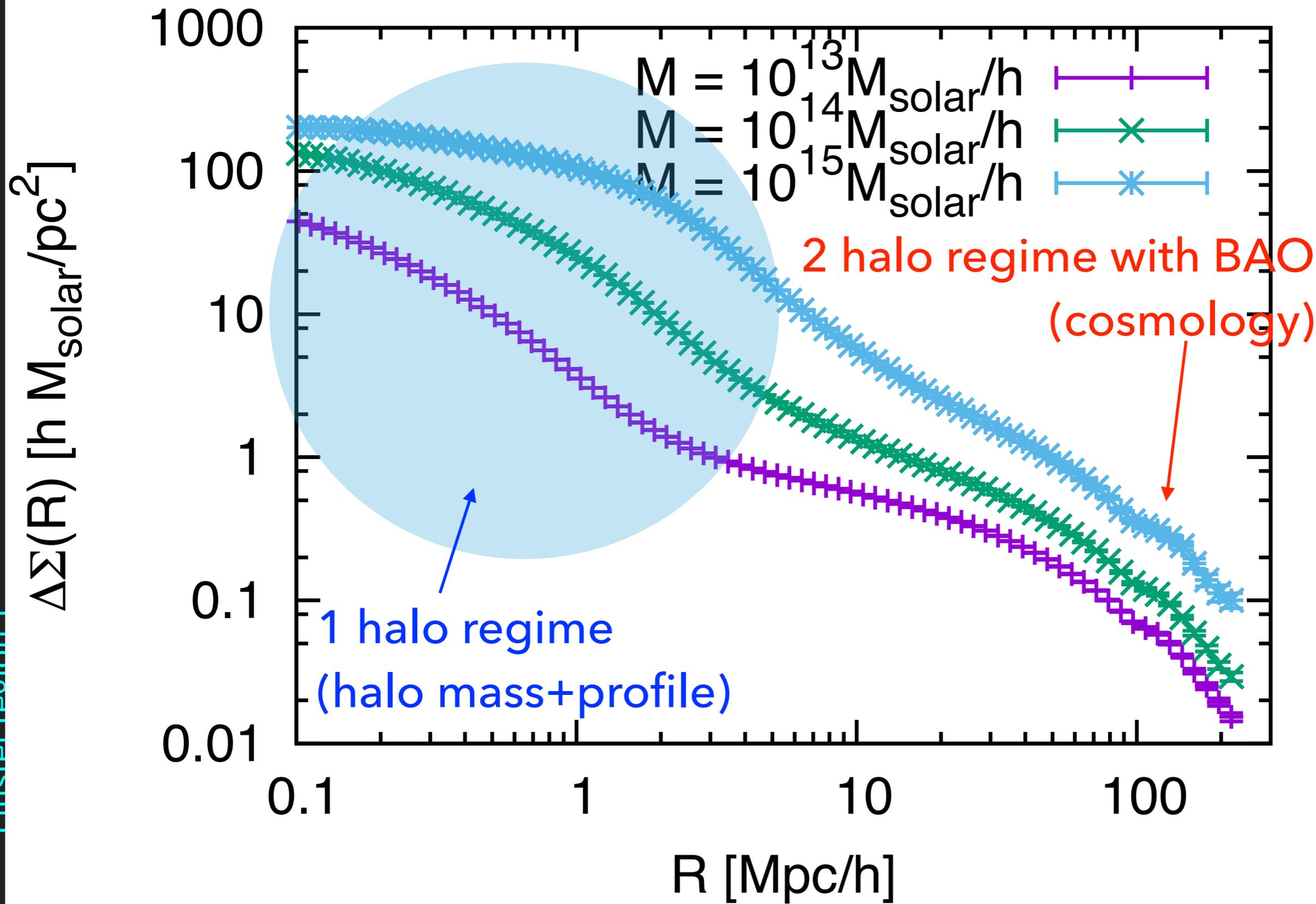
▶ **U. Tokyo**
Ken Osato*
Masamune Oguri

▶ **NAOJ**
Masato Shirasaki*
Takashi Hamana

▶ **Hirosaki U.**
Ryuichi Takahashi*



G-G LENSING SIMULATIONS



cluster region 1

da '11
i)
effect
 $\Sigma(R),$
 $[\bar{\Sigma}^2)] d\Pi$



(with z)

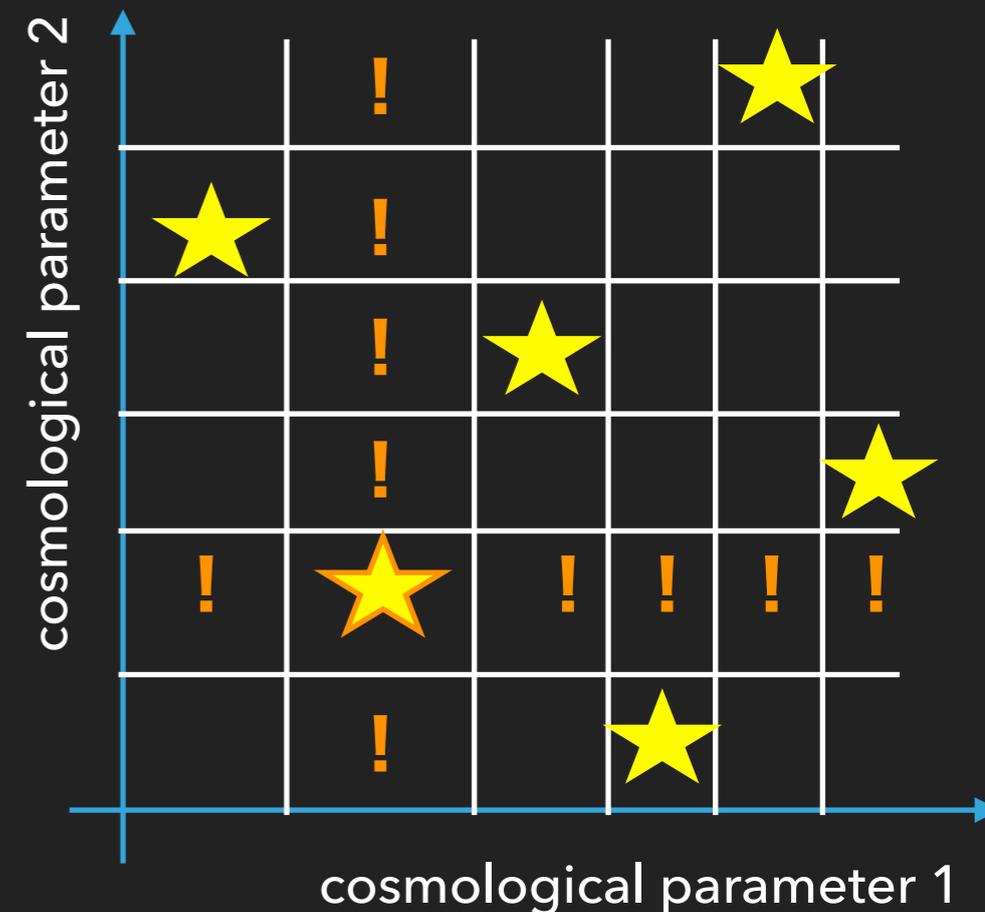
can probe the mean dark matter distribution around halos

SIMULATION SPECIFICATIONS

- ✓ N of particles: 2048^3
- ✓ box size: $1 h^{-1} \text{Gpc}$
resolve a $10^{12} h^{-1} M_{\text{solar}}$ halo with ~ 100 particles
- ✓ 2nd-order Lagrangian PT
initial condition @ $z_{\text{in}}=59$
(vary slightly for different cosmologies to keep the RMS displacement about 25% of the inter-particle separation)
- ✓ Tree-PM force by L-Gadget2
(w/ 4096^3 PM mesh)
 - ✓ ~ 2 days/run using 27 nodes on XC30 after parameter optimization
- ✓ 21 outputs in $0 \leq z \leq 1.5$
(equispaced in linear growth factor)
- ✓ Halo (subhalo) catalogs by
 - ✓ FOF + Subfind
 - ✓ Rockstar (+ merger tree by consistent-trees)
 - ✓ consistent central/satellite separation
- ✓ Data compression (256GB \rightarrow 48GB per snapshot)
 - ✓ positions \rightarrow displacement (16 bits per dimension; accuracy $\sim 1 h^{-1} \text{kpc}$)
 - ✓ velocity: discard after halo identification
 - ✓ ID: rearrange the order of particles by ID and then discard
 - ✓ already consuming $\sim 200 \text{TB}$ (100TB for DM, 100TB for halos) in half a year

EFFICIENT SAMPLING IN MULTI DIMENSIONAL SPACE: LATIN HYPERCUBE

- ✦ Each sample is the only one in each axis-aligned hyperplane containing it
 - ✦ One can find many realizations of such design (ex. diagonal design)
 - ✦ Impose additional condition such as “the sum of the distances to the nearest design point is maximal”
- ✦ **Useful technique to efficiently cover parameter space when the dimensionality is high but many simulations are not realistic**



EFFICIENT SAMPLING IN MULTI DIMENSIONAL SPACE: LATIN HYPERCUBE

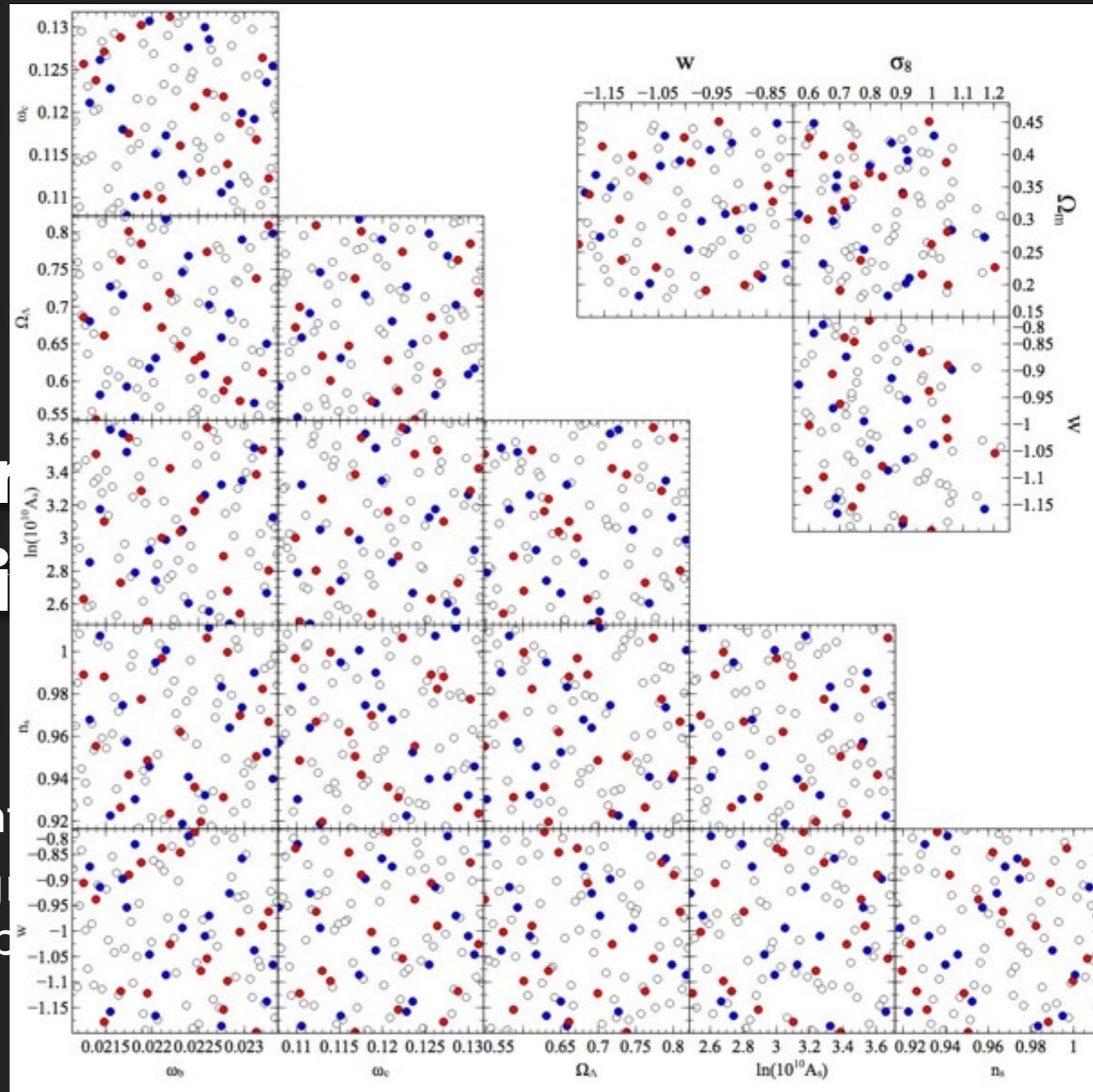
fiducial model

- ✓ PLANCK15 flat Λ CDM
- ✓ 24 realizations done
- ✓ assess statistical error
- ✓ check the accuracy of the emulator

84 sims
avail

varied cosmology

- ✓ "sliced" LH design (Ba, Brenneman & Myers '15)
- ✓ generate 100 samples even
- ✓ maxi-min distance LH design every 20 models (e.g., red/blue points)
- ✓ 2 types of sims



MEASURING G-G LENSING SIGNAL

- ✓ Work in 3D (c.f., 2D nature of lensing)
- ✓ And then projection on 2D
- ✓ **Hybrid Fourier-direct scheme**

3D cross spectrum in Fourier space

$$P_{\text{hm}}(\vec{k}) \text{ (on } 1024^3 \text{ mesh by FFT)}$$

Inverse FFT to real space and take the spherical average

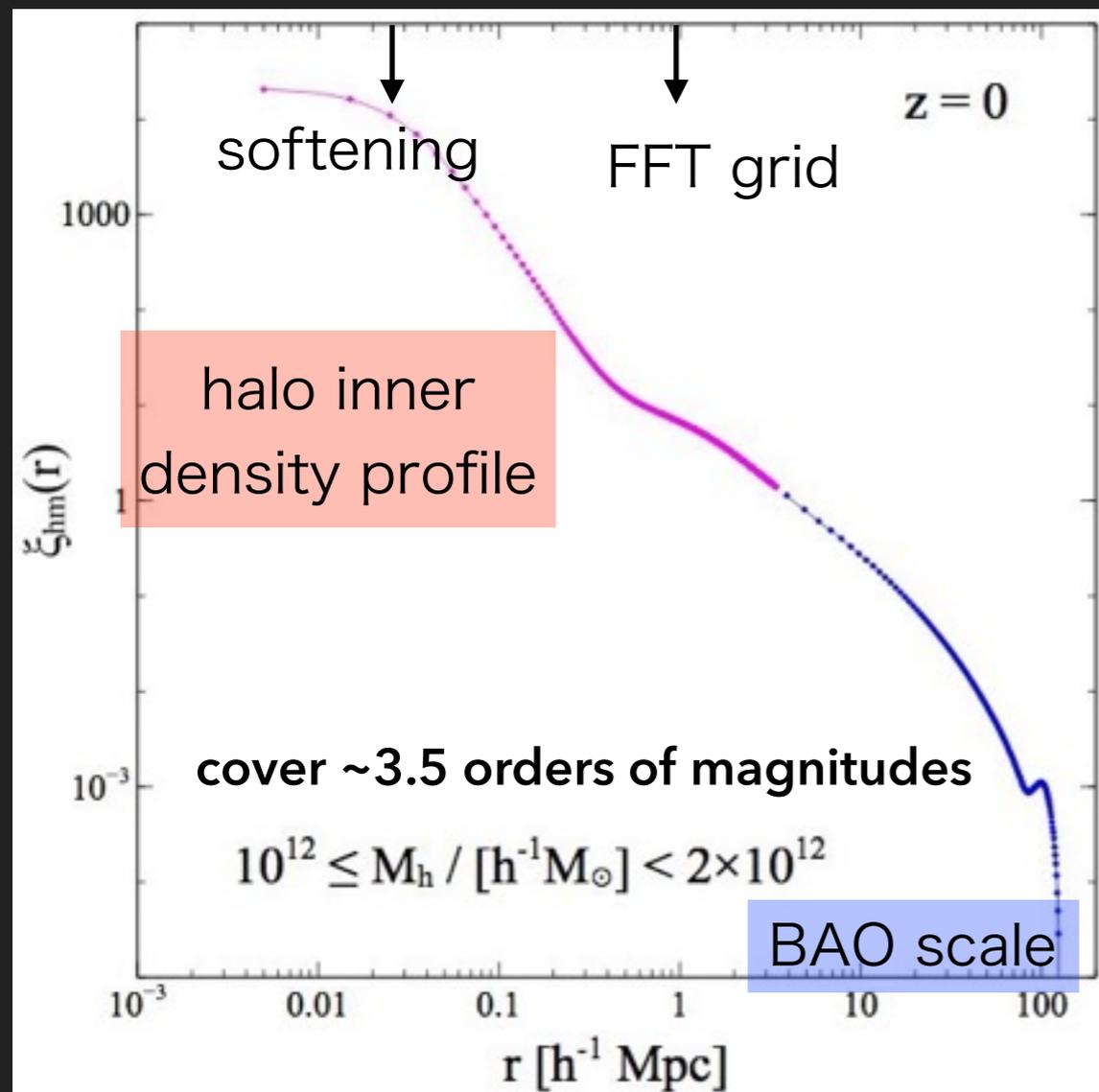
$$\xi_{\text{hm}}(\vec{r}) \xrightarrow{\text{spherical avg.}} \xi_{\text{hm}}(r)$$

ξ_{hm} on small scale from direct pair count

Finally project onto 2D to have $\Sigma(R)$ and then $\Delta\Sigma(R)$

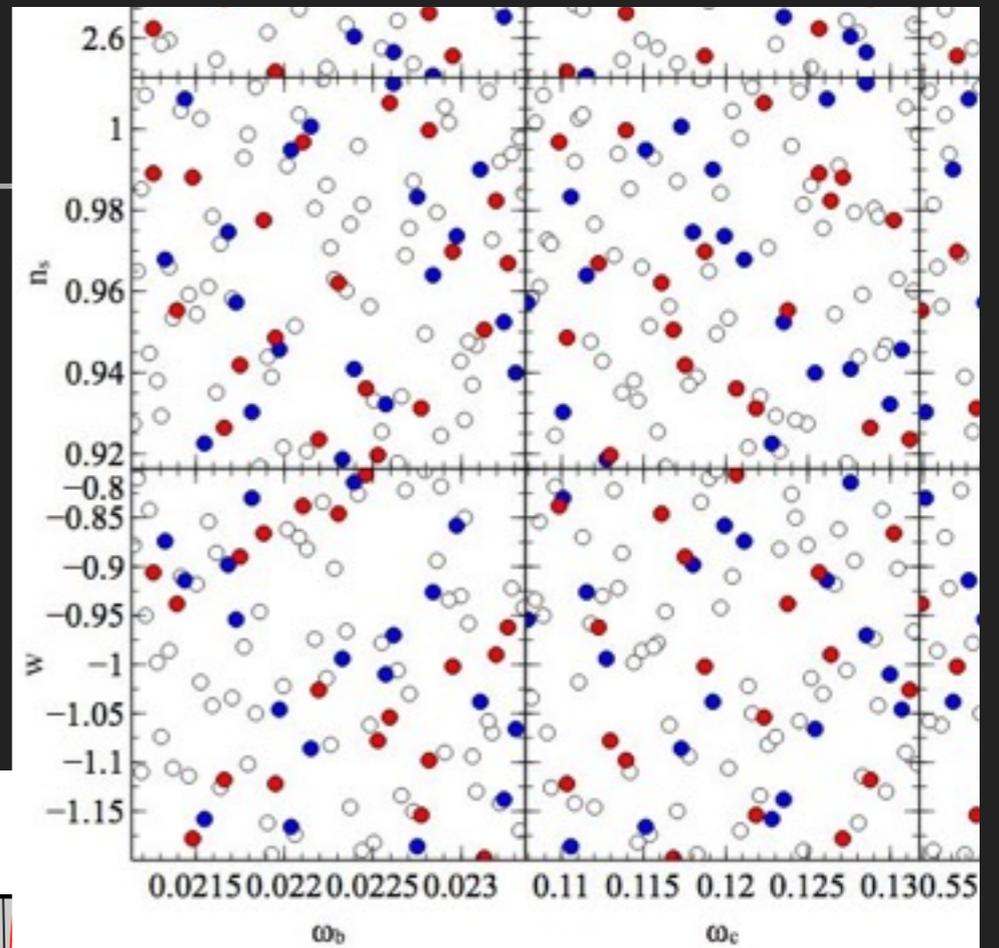
$$\Sigma(R) = \bar{\rho} \int [1 + \xi_{\text{gm}}(\sqrt{R^2 + \Pi^2})] d\Pi$$

Easier accuracy control than working in 2D, and quicker than direct pair counting alone



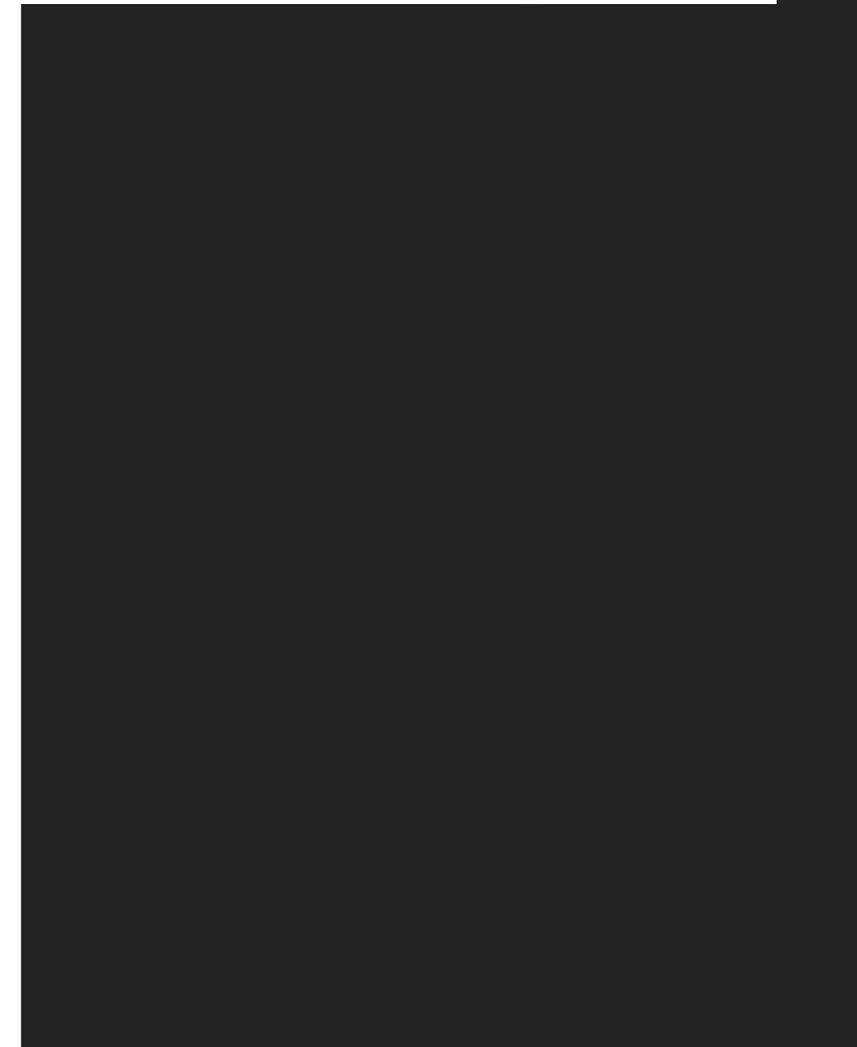
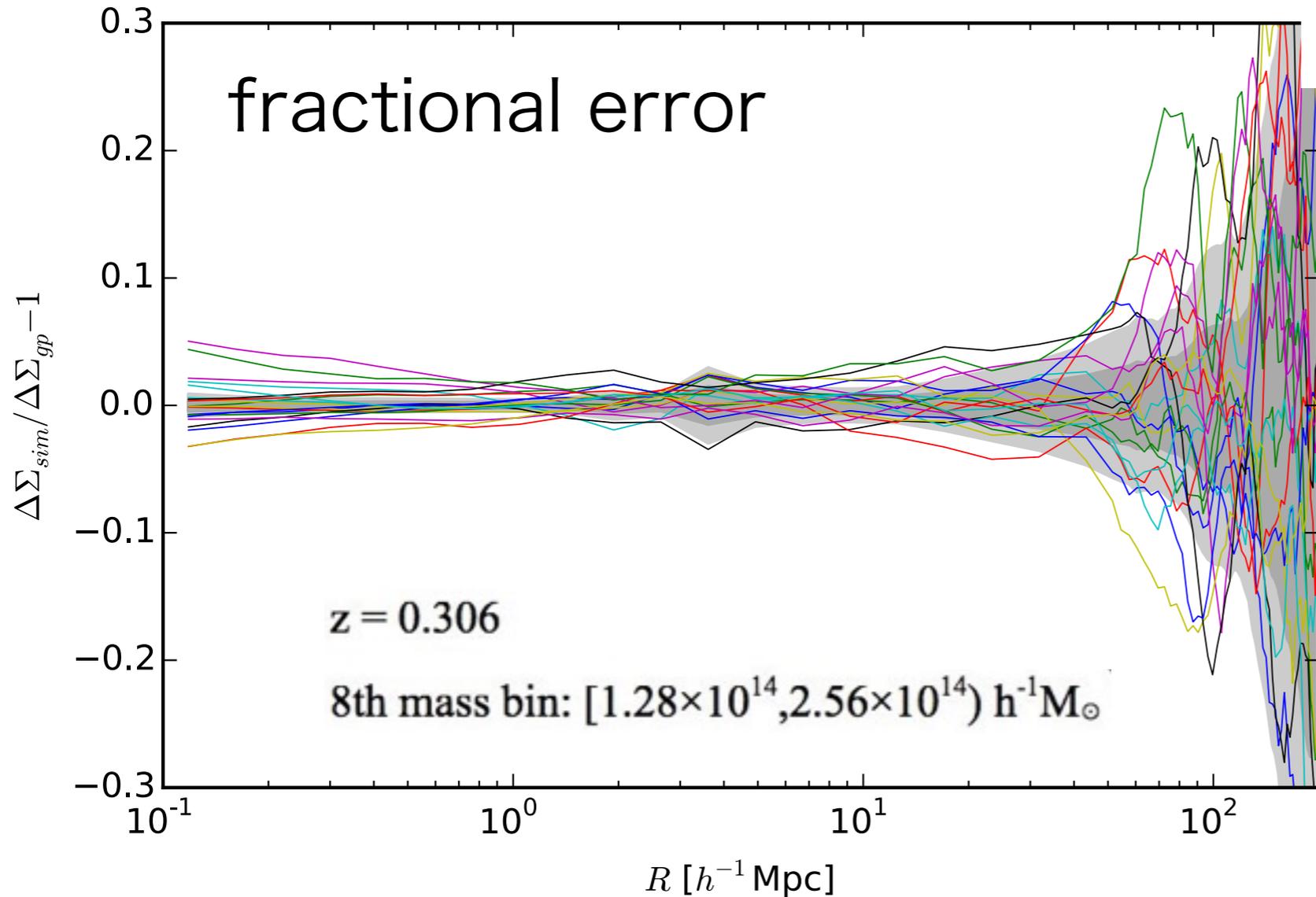
GAUSSIAN PROCESS

validation at 20 other models
(blue points)



random number varied

fractional error



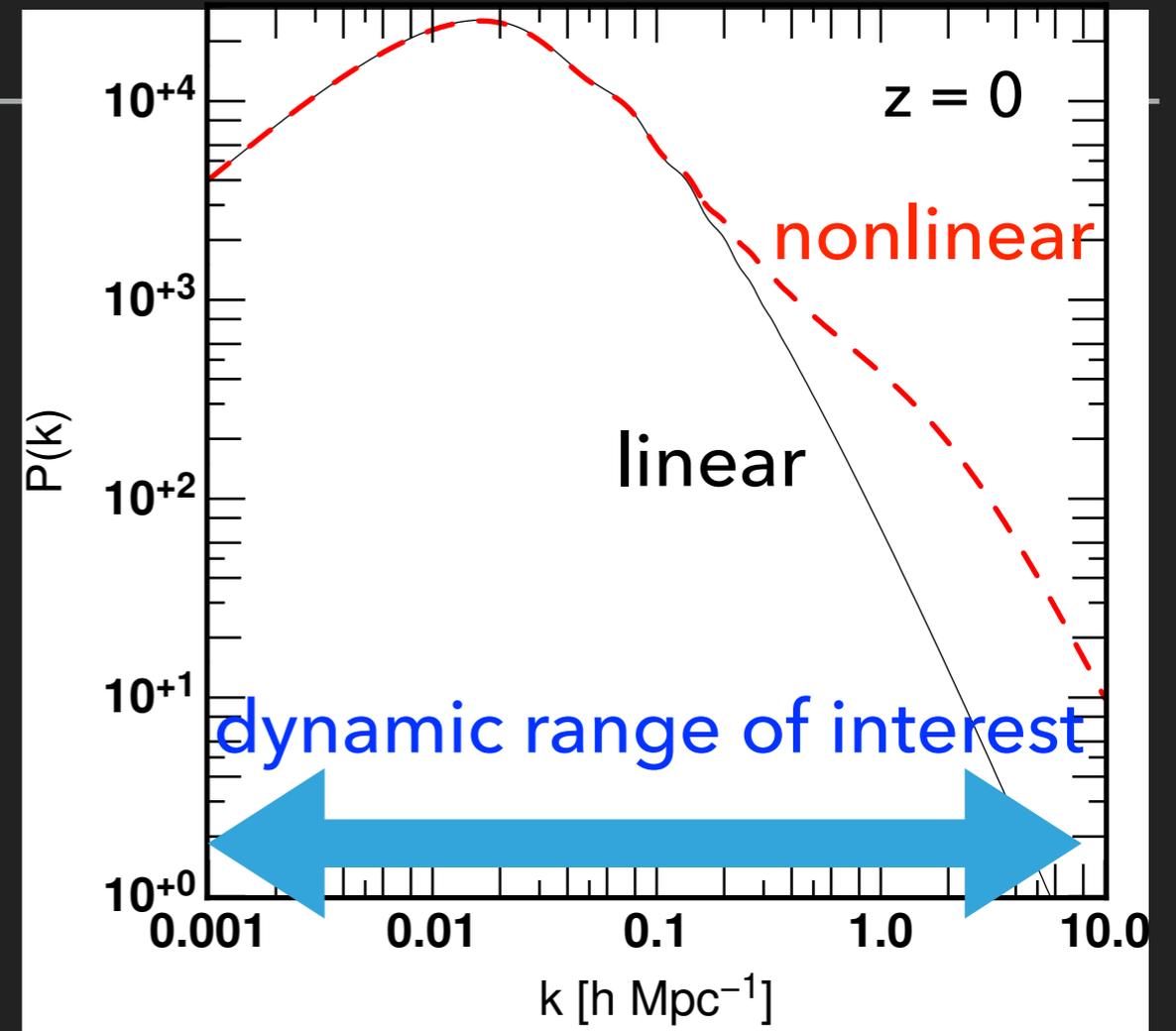
COMBINING PT WITH N-BODY

▶ Full-sky light cone simulation for future surveys (+ covariance)?

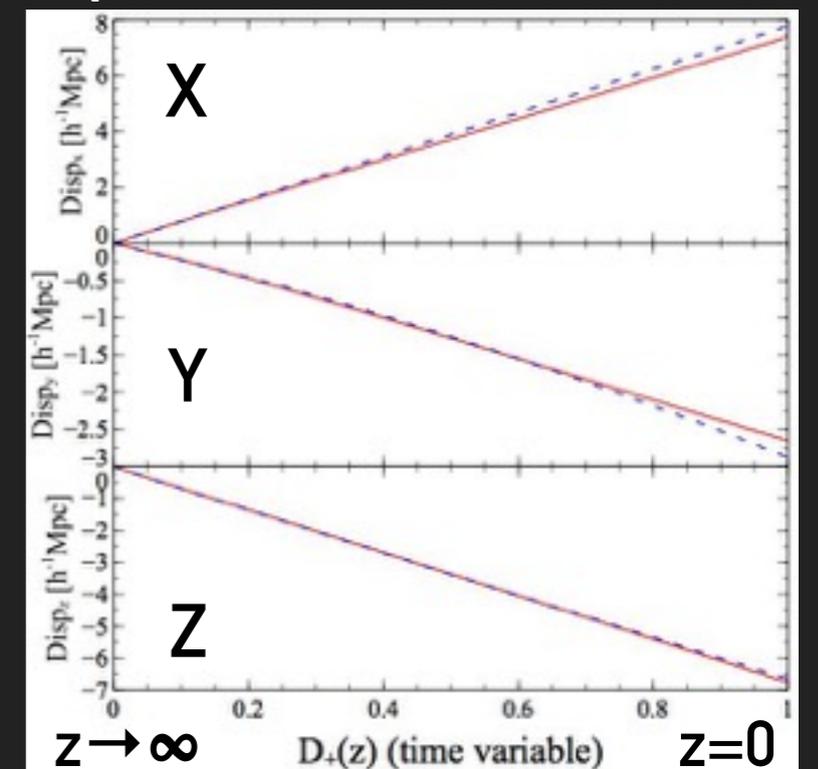
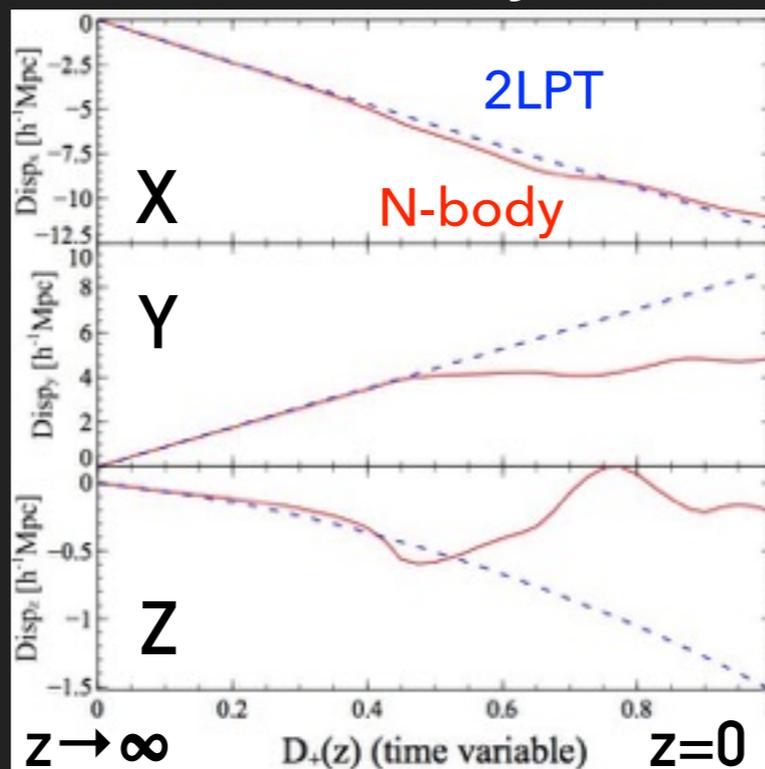
- ▶ Need a very large N to keep the resolution
- ▶ scalable?
- ▶ just solving linear dynamics on large scales

▶ Let us combine perturbative solution with N-body !

cosmic matter power spectrum



trajectories of particles



TREE-PM ALGORITHM w/ Naonori Sugiyama

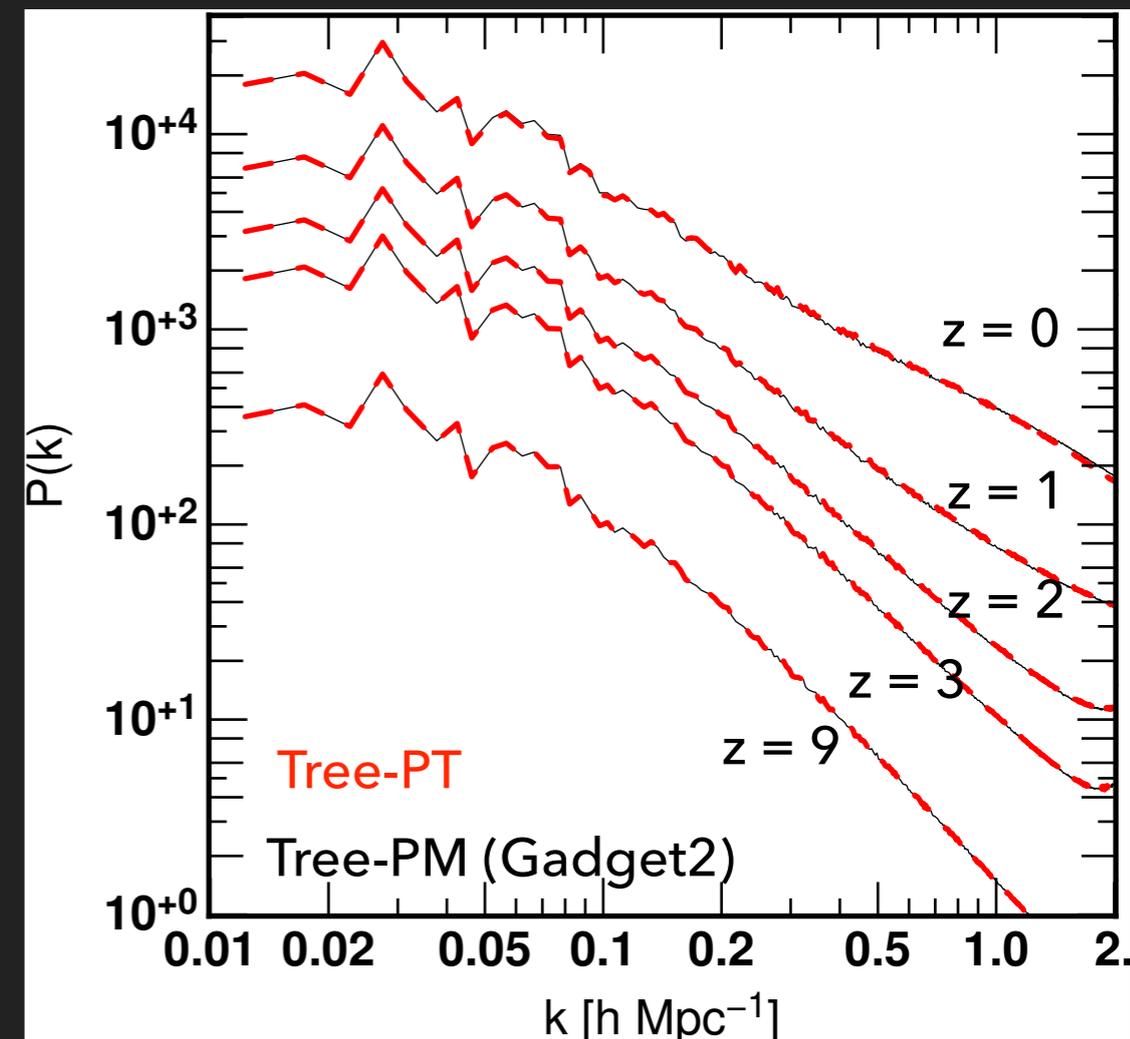
- ▶ Reinterpret PT solution as "PT force"

$$\vec{x}(\vec{q}; t) = \vec{q} + D_1(t)\vec{\Phi}^{(1)}(\vec{q}) + D_2(t)\vec{\Phi}^{(2)}(\vec{q})$$

- ▶ PT solution: displacement field
- ▶ Need to compute this just once
- ▶ "Force" can be defined as the time derivative of this
 - ▶ can be done analytically
- ▶ just solving linear dynamics on large scales
- ▶ Smooth out small scale contribution, replace the PM part in Gadget 2 to have Tree-PT
 - ▶ Still working on the "optimal" splitting scheme of the 2 forces
 - ▶ Speed is still an issue
- ▶ Developing a Tree-PM-PT code!

$L = 512 h^{-1}\text{Mpc}, N = 256^3$

$R_{\text{split}} = 10 h^{-1}\text{Mpc} \quad \sim 15\% \text{ speed up}$



SUMMARY

- ▶ **Modeling galaxy-galaxy lensing signal over a wide parameter/dynamic range**
 - ▶ a suite of cosmological simulations with Latin hypercube design
 - ▶ halo catalogs with density profile
 - ▶ Emulator based on Gaussian process (20 learning sets)
 - ▶ currently ~5% accuracy (20 validation sets) and hopefully this gets better
- ▶ **Quick simulation**
 - ▶ perturbative solution can be used as the long range force
 - ▶ Tree-PT code was developed based on Gadget2
 - ▶ Still need more work on the accuracy/speed optimization
 - ▶ Tree-PM-PT code under development